

1) F.W. , 8.012 Cl. Mech.

web page web.mit.edu/8.012/www

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recitations start next week

books : K+K. - primary text, only required. problems

BCG - more leisurely exposition of ~~except~~ basic material

alternative: How Things Work [www.wiley.com/college/howthingswork](http://www.wiley.com/college/howthingswork)

my "office hours"

questions during lecture / break

problem sets. due W at 9:30 AM - boxes { ~~it is posted~~  
due Sept. 18  
~~you shall get them back promptly~~  
answers will be posted right away  
you should get them back promptly

special problems... same procedure (note typo)

17 → 18

~~Holding~~ grading .3, .15, .15, .25, .15  
HW Q1 Q2 lab  
Oct. 10 Nov. 14

~~objections~~  
queries: in writing, ~~within 2 days~~ next to available recitation

policy on collaboration

lab → Rosenson

contact info.

Calendar

pb. sets + answers

sample exams; answers

"lecture notes"

①

2)

## scope and limit of classical mechanics

motion due to given forces.

some general rules about forces; some examples

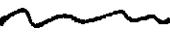
[other courses for details of what  
the fundamental forces are]

gravity  
~~spring~~  
 pulleys, ropes  
 ("')  
rigidity, constraint  
 friction

~~fantastic~~ extraordinary richness of phenomena

orbital mechanics: recent results  
 { sensitivity to initial conditions  
 Euler disc  
 Venus' spin

rattleback

break 

(3)

limits:

large velocities - special relativity  
 large distances } general relativity  
 very strong gravity }

fields, radiation

very small distances, structure of matter - quantum mechanics

BUT:

These builds on classical mechanics

$p, E, L$ , symmetry,  $m$

( $\times F!$ )

### 3) Elementary Concepts of CM

model of space:

triples of real numbers

( $\infty$  in all directions: locality)

Euclidean distance. homogeneous and isotropic

$$\xrightarrow{\text{to}} (x_1, x_2, x_3) \rightarrow \sqrt{(x_1 - y_1)^2 + (x_2 - y_2)^2 + (x_3 - y_3)^2}$$

model of time

real #s.

"intervals" only, (homogeneous)

## models of matter

point masses

→ position, velocity, acceleration

vectors ( $\vec{v}$  vs. speed)

→ vector, after fixing an origin

build up extended bodies from pt.!

→ intrinsic property. positive #  $\times$  unit (gm., kg.)

we'll discuss its measurement + use next time

not weight : astrophysics

(vs. supermarket)

0<sup>th</sup> law: conservation of mass - fr pt. mass, no change  
build up + break apart bodies!

Central Thought:

4) ~~The Primacy of Acceleration  
(not velocity)~~ (Newton's 1<sup>st</sup> law)

{ bicyc

4) Newton's 1<sup>st</sup> Law: "Natural Motion"

constant velocity

{ bicyclists  
cars  
ice-skating  
astronauts

Galileo's ship ; Earth in space! (rotation + revolution)

vs. { friction  
bacteria

Sept. 5

5) Newton's 2<sup>nd</sup> Law

$$\vec{F} = m\vec{a} = m\frac{d\vec{v}}{dt} \quad (\vec{F} = \frac{d}{dt}(m\vec{v}))$$

empty? (defining force)

character of forces

must have a nearby source

simple  
dependent  
on mass

gravity: force of  $m$  universal  
(so just  $\vec{a}$ !)

all others\*: force independent of  $m$

walking  
chain  
demo.

## 6] Examples of forces

near-Earth gravity

$$\vec{F} = m\vec{g} \quad (= m\vec{a})$$

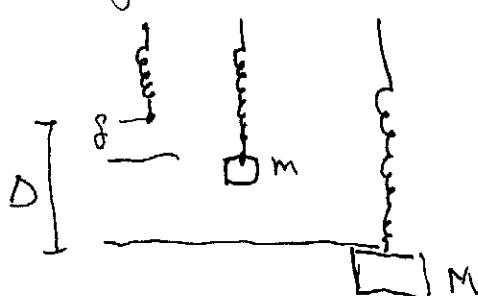
$$\vec{g} = 32 \text{ ft/sec}^2 \text{ or } 980 \text{ cm/sec}^2 \text{ down}$$

Spring

$$\vec{F} = -k\vec{x} \quad \vec{x} = \text{displacement from equilibrium}$$

(Hooke's law)  $k$  = spring constant

"weighing" mass



Free diagram:

$$\begin{matrix} \uparrow k\delta \\ \downarrow mg \end{matrix}$$

$$m = \frac{k\delta}{g}; \quad M = \frac{k\Delta}{g},$$

$$\frac{m}{M} = \frac{\delta}{\Delta}$$

other ways of to weigh: balance, collisions  
they must all turn out consistent!

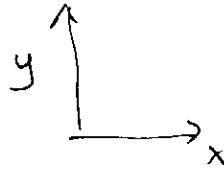
## 7] Examples of motion

### a) constant acceleration (projectiles)

$$\ddot{x} = 0$$

(functions  
of  $t$ )

$$\ddot{y} = -g$$

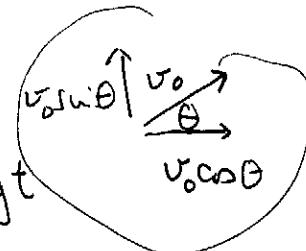


$$\dot{x} = c_1, \quad \dot{y} = c_2 - gt$$

$$= v_{0x} \quad = v_{0y} - gt$$

$$x = x_0 + v_{0x} t$$

$$y = y_0 + v_{0y} t - \frac{1}{2} g t^2$$



Problem: given magnitude of initial velocity,  
determine angle for maximum range

Sol<sup>n</sup>: choose  $x_0 = y_0 = 0$ ,  $v_{0x} = v_0 \cos \theta$ ,  $v_{0y} = v_0 \sin \theta$

$$x = v_0 \cos \theta t$$

$$y = v_0 \sin \theta t - \frac{1}{2} g t^2$$

$$\text{range} = v_0 \cos \theta t_f$$

$$0 = v_0 \sin \theta t_f - \frac{1}{2} g t_f^2 \Rightarrow t_f = \frac{2 v_0 \sin \theta}{g}$$

$$\text{range} = \frac{2 v_0^2}{g} \sin \theta \cos \theta = \frac{v_0^2}{g} \sin 2\theta \Rightarrow \theta = \pi/4 \text{ for max.}$$

7

b) uniform circular motion



$$\theta = \omega t$$

$$x = R \cos \omega t$$

$$y = R \sin \omega t$$

$$\dot{x} = -R \omega \sin \omega t$$

$$\dot{y} = R \omega \cos \omega t$$

$$\ddot{x} = -R \omega^2 \cos \omega t = -\omega^2 x$$

$$\ddot{y} = -R \omega^2 \sin \omega t = -\omega^2 y$$

requires force pointing radially inward  
in break

8) Brief mathematical interlude: vectors

quantities with magnitude and direction

examples: position relative to a chosen origin

velocity

displacement

velocity

acceleration

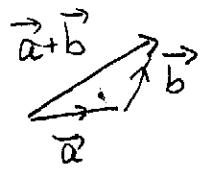
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different units!

After choosing axes, a vector can be written as a triple of numbers (with units) - the coordinates of its end-point

Notations:  $\vec{v}$ ;  $(v_1, v_2, v_3)$ ;  $\begin{pmatrix} v_1 \\ v_2 \\ v_3 \end{pmatrix}$ ;  $(v_x, v_y, v_z)$ ;  
 $\mathbf{v}$  (boldface),  $v$

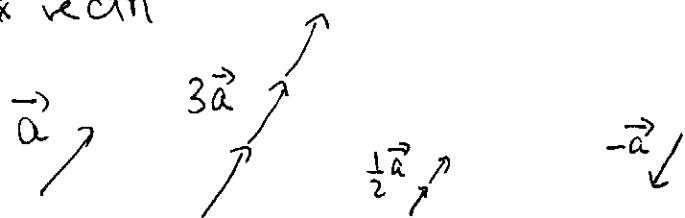
## vector addition



Newton 2.

addendum to Newton 2: forces from different sources add as vectors!

scalar \* vector



$$\vec{a} + (-\vec{a}) = 0$$

$$(\vec{a} + \vec{b}) + \vec{c} = \vec{a} + (\vec{b} + \vec{c})$$

$$\lambda(\vec{a} + \vec{b}) = \lambda\vec{a} + \lambda\vec{b}$$

If written as triples, just {add (each piece independently)}  
{multiply}

interpretation of components:

$$\vec{v} = v_1 \vec{i} + v_2 \vec{j} + v_3 \vec{k}$$

$$(or \quad v_1 \hat{x} + v_2 \hat{y} + v_3 \hat{z})$$

unit vector in  $\vec{x}$  direction

magnitude

$$|\vec{v}| = \sqrt{v_1^2 + v_2^2 + v_3^2} \quad \text{upt. 10}$$

you can take derivatives of vects that depend on time, to get other vects.

dot product

$$\vec{v} \cdot \vec{\omega} = v_1 \omega_1 + v_2 \omega_2 + v_3 \omega_3$$

~~see any  
text!~~

$$= |\vec{v}| |\vec{\omega}| \cos \angle(\vec{v}, \vec{\omega})$$

$$|\vec{v} + \vec{\omega}|^2 = (\vec{v} + \vec{\omega}) \cdot (\vec{v} + \vec{\omega})$$

$$\vec{v} \cdot \vec{\omega} = \vec{\omega} \cdot \vec{v}$$

$$= |\vec{v}|^2 + |\vec{\omega}|^2 + 2 \vec{v} \cdot \vec{\omega}$$

$$\vec{v} \cdot (\vec{v} + \vec{\omega}) > \vec{v} \cdot \vec{\omega} + \vec{\omega} \cdot \vec{v}$$

"law of cosines"

$$\vec{v} \cdot \vec{\omega} = |\vec{v}| |\vec{\omega}| \cos \angle(\vec{v}, \vec{\omega})$$

$$\vec{v} \cdot \vec{\omega} = 0 \iff \vec{v} \perp \vec{\omega}$$

example: circular motion

$$\vec{r}(t) = (R \cos \omega t, R \sin \omega t, 0)$$

$$\vec{v}(t) = (-R \omega \sin \omega t, R \omega \cos \omega t, 0) \quad |\vec{v}| = R \omega$$

$$\vec{a}(t) = (-R \omega^2 \cos \omega t, -R \omega^2 \sin \omega t, 0) \quad |\vec{a}| = R \omega^2$$

$$\vec{r} \cdot \vec{v} = 0 ; \quad \vec{a} = -\omega^2 \vec{r}$$

$$\text{why? } \vec{r}^2 = R^2 \Rightarrow 2 \vec{r} \cdot \frac{d\vec{r}}{dt} = 0 \quad \rightarrow \vec{r}^2 = R^2 \Rightarrow \vec{r} \cdot \vec{a} = 0 \quad \uparrow \text{!}$$

## 9) Inertial and non-inertial frames

(to "in practice" experiments)

Postulate: It is possible to set up observers with identical clocks, rulers, have them "stand still" and communicate instantaneously (to synchronize watches). Their measurements will conform to Newton's laws. Inertial frame

If you happen not to be working in an inertial frame, there are extra "fictitious" forces



$$mg_{\text{grav.}} - kx = m\omega^2 R$$

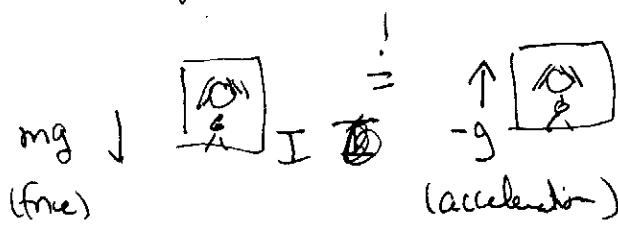
$$m(g_{\text{grav.}} - \omega^2 R) = kx$$

Geff. "

Note "fictitious" forces are  $\propto m$  (that was the \*)

10) Einstein: ~~relativity~~ "happiest thought of my life"

effect of gravity equivalent to accelerated frame



$$\text{I: } m\vec{a}_{\text{observed}} = \vec{mg} ; \vec{a}_{\text{observed}} = \vec{g}$$

$$\text{II: } m\vec{a}_{\text{total}} = 0 \quad \vec{a}_{\text{total}} = \vec{a}_{\text{observed}} + (-\vec{g}) ; \vec{a}_{\text{observed}} = \vec{g}$$

(11)

Advantage of II: explains why  $\vec{F} \propto m!$

## 10] Mach's Principle

Consider an isolated whirling pair

Dense  $\square$   
 $\downarrow$   
 $\uparrow$

vs. stationary  $\square$  less  $\square$

From the length of the spring you can tell ~~if~~ who's  
 they're rotating (accelerating).

Mach: Accelerated relative to what?

fixed stars! ?? - not in existing physics  
 ~ 17 Sept.

## 11] Units and Dimensional Analysis

$L, T, M$   
 in any valid eq<sup>n</sup>, ~~dimensions~~<sup>"dimensions"</sup> must match (choice of  
 units can't matter)

$$\text{e.g. } [a] = L/T^2$$

$$[s] = L$$

$$\text{uniform motion: } \cancel{s} = L$$

$$s = f(a, t) \xrightarrow{x \propto a} \xrightarrow{x \propto t}$$

change unit of length, time

$$\lambda s = f\left(\frac{\lambda}{a} a, \varepsilon t\right)$$

(11a)

uniform motion with initial velocity  
 $(s = vt - \frac{1}{2} at^2)$

$$\lambda s = f\left(\frac{\lambda}{\varepsilon^2} a, \frac{\lambda}{\varepsilon} v, \varepsilon t\right)$$

$$\varepsilon = 1/t, \lambda = 1/at^2$$

$$\frac{s}{at^2} = f(1, \frac{v}{at}, 1)$$

$$\text{in fact, } \frac{s}{at^2} = \boxed{\frac{v}{at}} - \frac{1}{2}$$

(to 12)

(12)

$$\text{choose } \varepsilon = 1/t, \quad \lambda = \frac{\varepsilon^2}{a} = \frac{1}{at^2}$$

$$\frac{S}{at^2} = f(1, 1) = \text{Froude Number} \\ (\text{to 1/a})$$

uniform circular motion

$$[F] = ML/T^2$$

$$[m] = M, [R] = L, [\omega] = 1/T$$

$$F = f(m, R, \omega)$$

$$F \frac{\lambda m}{\varepsilon^2} = f(\lambda m, mR, \omega/\varepsilon)$$

$$\text{with } \lambda = \frac{1}{n}, m = \frac{1}{R}, \varepsilon = \omega$$

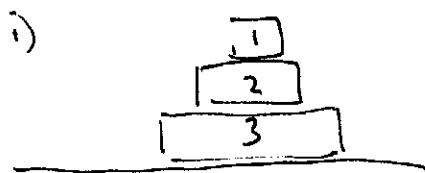
$$\frac{F \cancel{\varepsilon}}{mR\omega^2} = f(1, 1, 1) = \text{number} \\ \leftarrow \text{wind tunnels, ...} \sim \tilde{J}^{9/12}$$

## 12) Newton's 3rd Law

$$\xrightarrow{\text{Action}} = (-) \xrightarrow{\text{reaction}}$$

deep applications: cons. of momentum  
extended bodies

example 1:



$$\text{Forces on 1: } \vec{F}_{12} + m_1 \vec{g} = 0$$

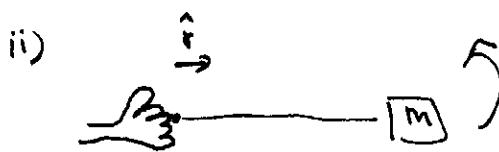
$$\text{on 2: } \vec{F}_{21} + \vec{F}_{23} + m_2 \vec{g} = 0$$

$$\text{on 3: } \vec{F}_{3f} + \vec{F}_{32} + m_3 \vec{g} = 0$$

$$\text{on floor: } \vec{F}_{f3} \quad (= M \ddot{\vec{a}}) \\ \uparrow \text{very big!}$$

$$\begin{aligned} \vec{F}_{f3} &= -\vec{F}_{3f} = \vec{F}_{32} + m_3 \vec{g} \\ &= -\vec{F}_{23} + m_3 \vec{g} \\ &= \vec{F}_{21} + m_2 \vec{g} + m_3 \vec{g} \\ &= -\vec{F}_{12} + m_2 \vec{g} + m_3 \vec{g} \\ &= m_1 \vec{g} + m_2 \vec{g} + m_3 \vec{g} = (m_1 + m_2 + m_3) \vec{g} \\ &= \text{W.I.S.B.} \end{aligned}$$

(What It Should Be!)



$$\vec{F}_{\text{block, rope}} = -m R c \omega^2 \hat{r}$$

$$\vec{F}_{\text{rope, hand}} + \vec{F}_{\text{rope, block}} \approx m_{\text{rope}} \approx 0 \quad (\text{Simple case})$$

$$\vec{F}_{\text{hand, rope}} = -\vec{F}_{\text{rope, hand}} \approx \vec{F}_{\text{rope, block}} = -\vec{F}_{\text{rope, block}} = m R c \omega^2 \hat{r}$$

"centrifugal force"

### 3) Constraint forces



Bodies supply their "automatically" to keep  
matter from penetrating  
(Pauli exclusion principle)

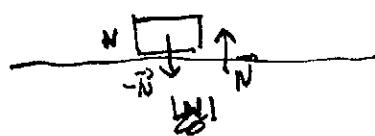
### 4) Friction forces

very difficult to explain from 1<sup>st</sup> principles

(also a very important practical subject)

- lubrication : ball bearings
- + salting : roughening of tracks, tracks

frequently useful phenomenological  
description:



static:  $\vec{F}_{fr.}$  tangential

$$|\vec{F}| \leq \mu_s |\vec{N}|$$

dynamics

$$\vec{F}_{fr.} = \mu_d |\vec{N}|, \text{ directed against}$$

the motion ( $\Rightarrow \vec{F}_{fr.} \cdot \vec{v}_{rel.} < 0$ )

$$\vec{F}_{fr.} = -\hat{v}_{rel.} \mu_d |\vec{N}|$$

$$= -\frac{\hat{v}_{rel.}}{|\vec{v}_{rel.}|} \mu_d |\vec{N}|$$

$$\vec{F}_{fr.} = -\hat{v}_{real.} \mu_s |\vec{N}|$$

$\sim T_{9/2}$

example: rolling without slipping (train)



fixed : static friction

It's static friction that moves trains!